# Echo of the quantum phase transition of CeCu<sub>6-x</sub>Au<sub>x</sub> in XPS: Breakdown of Kondo screening

M. Klein,<sup>1</sup> J. Kroha,<sup>2</sup> H. v. Löhneysen,<sup>3,4</sup> O. Stockert,<sup>5</sup> and F. Reinert<sup>1,6</sup>

<sup>1</sup>Experimentelle Physik II, Universität Würzburg, Am Hubland, D-97074 Würzburg, Germany

<sup>2</sup>Physikalisches Institut and Bethe Center for Theoretical Physics, Universität Bonn, Nussallee 12, D-53115 Bonn, Germany

<sup>3</sup>Physikalisches Institut, Universität Karlsruhe (TH), D-76128 Karlsruhe, Germany

<sup>4</sup>Forschungszentrum Karlsruhe, Institut für Festköperphysik, D-76021 Karlsruhe, Germany

<sup>5</sup>Max Planck Institute for Chemical Physics of Solids, Nöthnitzer Strasse 40, 01187 Dresden, Germany

<sup>6</sup>Forschungszentrum Karlsruhe, Gemeinschaftslabor für Nanoanalytik, D-76021 Karlsruhe, Germany

(Received 5 November 2008; published 12 February 2009)

We present an x-ray photoemission study of the heavy fermion system  $\text{CeCu}_{6-x}\text{Au}_x$  across the magnetic quantum phase transition of this compound at temperatures above the single-ion Kondo temperature  $T_K$ . In dependence of the Au concentration x we observe a sudden change in the *f*-occupation number  $n_f$  and the core-hole potential  $U_{df}$  at the critical concentration  $x_c=0.1$ . We interpret these findings in the framework of the single-impurity Anderson model. Our results are in excellent agreement with findings from earlier ultraviolet photoemission spectroscopy measurements and provide further information about the precursors of quantum criticality at elevated temperatures.

DOI: 10.1103/PhysRevB.79.075111

PACS number(s): 71.27.+a, 71.28.+d, 79.60.-i, 71.10.-w

#### I. INTRODUCTION

 $\rm CeCu_{6-x}Au_x$  is one of the best characterized heavy-fermion (HF) compounds^{1-5} exhibiting a quantum phase transition (QPT) between a paramagnetic phase with full Kondo screening of the Ce 4f moments and an antiferromagnetic phase induced by their Ruderman-Kittel-Kasuya-Yoshida (RKKY) coupling.<sup>6</sup> The QPT occurs at a critical Au concentration of  $x=x_c=0.1$ . Two major scenarios for the QPT have been put forward, first the Hertz-Millis scenario<sup>7,8</sup> where only the (bosonic) magnetic fluctuations become critical, leaving the (fermionic) heavy quasiparticles intact, and second the local quantum criticality<sup>9,10</sup> where the local Kondo quasiparticles are destroyed by coupling to the quantum critical magnetic fluctuations. In the latter case the Kondo screening scale or Kondo temperature of the lattice system,  $T_{K}^{*}$ , vanishes at the quantum critical point (QCP). Despite intense efforts it has not been possible to unambiguously identify either one scenario in CeCu<sub>6-x</sub>Au<sub>x</sub>. The difficulty resides in the fact that near the QCP many different effects come into play, including local Kondo screening, lattice coherence and Fermi volume collapse, quantum critical fluctuations, and possible dimensional reduction.<sup>2,11</sup> They have prevented a conclusive experimental picture as well as a unified theory of the QPT. Recently we have presented direct measurements of the evolution of the local Kondo screening energy scale  $T_K$  across the QCP as extracted from high-resolution ultraviolet photoemission spectroscopy (UPS).<sup>12</sup> Taking impurity spectra clearly above  $T_K$  as well as above the lattice coherence temperature  $T_{\rm coh}$  and the Néel temperature for magnetic ordering,  $T_N$  ( $T > \{T_K, T_{coh}, T_N\}$ ), made it possible to probe the local energy scale without the complications caused by lattice coherence or quantum critical fluctuations. The surprising outcome of these investigations is that the precursors of the QPT are already visible in the single-impurity  $T_K$  as extracted from UPS spectra at such elevated temperatures. In Ref. 12 we developed a theory based on the effective single-impurity model realized for temperatures  $T > T_{\rm coh}$ , where the Kondo exchange coupling J of a Ce site is self-consistently renormalized by the surrounding identical Ce atoms only through the indirect RKKY coupling. The theory explains the observed nontrivial step-like behavior of  $T_K$  near the QCP and supports the local critical scenario for CeCu<sub>6-x</sub>Au<sub>x</sub>. It also provides a general criterion to distinguish experimentally whether a given HF compound should exhibit Hertz-Millis (HM) or local quantum critical behavior.<sup>12</sup>

UPS provides the most direct access to the screening scale  $T_K$  by directly recording the Kondo resonance in the local Ce 4f spectrum<sup>13,14</sup> but probes a relatively shallow surface region only. Therefore we complement here the studies of Ref. 12 with bulk-sensitive x-ray photoemission spectroscopy (XPS) on the same samples and under the same experimental conditions even though XPS provides only rather indirect information about the Ce 4f spectrum and the physics near the Fermi energy  $E_F$ .<sup>15,16</sup> We find that the XPS results are in full agreement with the UPS analysis.

### **II. SIGNIFICANCE OF CORE-HOLE SPECTRA**

Before we show the experimental results, we briefly recall the energetics of the photoionization process, in order to understand the essential structure of the XPS spectra. In XPS the Ce 3*d* spectrum is recorded in the final state of the Ce ion after creation of a 3*d* core hole. The binding energy of a 3*d* electron is influenced by the electrostatic attraction  $-U_{df}$  between the 3*d* core hole and a 4*f* electron and, hence, depends on the number of electrons *n* in the 4*f* shell. Therefore, the 3*d* peak of the XPS spectrum is split into individual resonances, each one corresponding to a different charge state of the 4*f* shell, the so-called  $f^n$  resonances, with n=0,1,2. In addition, there is a Coulomb repulsion  $U_{ff}$  between two electrons in the 4*f* shell. The peak position (binding energy  $E_B^{(n)}$ ) of the  $f^n$  resonance is the energy difference between the charge configuration in the photoionized final state and in the



FIG. 1. (Color online) 3d core-level spectra of  $\text{CeCu}_{6-x}\text{Au}_x$  for five different Au concentrations at T=15 K (x-ray energy  $h\nu$ =1486 eV and experimental resolution  $\Delta E \approx 2$  eV). The vertical lines mark the peak positions exhibiting an abrupt shift of the  $f^1$  and  $f^2$  resonances from x=0.1 to x=0.2.

ground state. Hence, it is roughly given by, up to conductionelectron screening corrections,

$$E_B^{(n)} = E_B^{(0)} - nU_{df} + \frac{1}{2}(n^2 - n)U_{ff}.$$
 (1)

Note that the attractive core-hole potential  $-U_{df}$  changes the whole level scheme.<sup>17</sup> In particular, in the photoionized state the  $4f^2$  configuration becomes occupied (see Fig. 1) even though in the Ce ground state it is shifted far above the Fermi energy by the repulsion  $U_{ff}$ . The quantitative fit of  $U_{df}$  and  $U_{ff}$  (see below) shows that the binding energy of the  $4f^0$  configuration,  $E_B^{(0)}$ , is the largest. The spectral weight of the  $f^0$  XPS peak is proportional to the Kondo resonance of the Ce 4f spectrum since in CeCu<sub>6-x</sub>Au<sub>x</sub> most of the Kondo peak weight is located above  $E_F$  (Ref. 13) and thus represents the unoccupied configuration of the 4f shell.

#### **III. EXPERIMENTAL RESULTS**

All the measurements were carried out at T=15 K and thus at temperatures far above any long-range or quantum critical fluctuations. The experimental setup is equipped with a SCIENTA R4000 analyzer and a monochromatized x-ray source (Al K<sub> $\alpha$ </sub> line: x-ray energy  $h\nu$ =1486 eV). The pressure in the UHV chamber was  $1 \times 10^{-10}$  mbar. Because of the low intensity due to the small surface area of the crystals and the rapid deterioration of the surface, the instrumental resolution was set to  $\approx 2$  eV in order to optimize the count rate. The investigated samples were single crystals grown in a W crucible under high-purity Ar atmosphere<sup>18</sup> and cleaved *in situ* at the sample position just before the measurement.

The Ce 3d core-level spectra for all measured compounds are shown in Fig. 1. Taking the second derivative of each spectrum reveals that each of the broad peaks is actually comprised of two resonances. Hence, each spectrum consists of a group of three resonances, termed  $f^0$ ,  $f^1$ , and  $f^2$  in Fig. 1. They are, in addition, duplicated due to a spin-orbit splitting of  $\Delta_{SO}=18.9$  eV. A rough analysis already shows that the XPS 3*d* spectra can be divided into two classes, namely, for  $x \le 0.1$  and  $x \ge 0.2$ , respectively, such as in the case of the UPS 4*f* spectra:<sup>12</sup> for  $x \le 0.1$  the  $f^0$  peak at  $E_B \approx 917$  eV is clearly visible, albeit comparatively weak, while for  $x \ge 0.2$ the  $f^0$  weight is almost not discernible and vanishes completely for x=1. In addition, for  $x \ge 0.2$  the  $f^1$  and  $f^2$  peaks are shifted by a constant amount of 1.8 eV toward smaller binding energy relative to their positions for  $x \le 0.1$ .

#### **IV. DISCUSSION**

In order to interpret these XPS results quantitatively, we calculated the Ce 3d spectra using the Gunnarsson-Schönhammer (GS) theory,<sup>19,20</sup> which is based on the singleimpurity Anderson model (SIAM). From earlier UPS experiments it is known<sup>12,13,21</sup> that the CeCu<sub>6-x</sub>Au<sub>x</sub> 4f spectra and their T dependence are very well described in the framework of this model for elevated temperatures  $T > T_K$ , with an RKKY-induced renormalization of the local Kondo coupling.<sup>12</sup> Moreover, the GS theory has been successfully applied to interpret XPS, x-ray photoabsorption, inverse photoemission, and valence photoemission spectroscopy on other Ce-based compounds, using the same parameter set for all measurements,<sup>22,23</sup> albeit small adjustments of parameter values were pointed out to be necessary in some cases.<sup>24,25</sup> In our GS calculations for CeCu<sub>6-r</sub>Au<sub>r</sub> we use, without fitting, the same parameter values of the SIAM that were determined earlier from noncrossing approximation (NCA) calculations of the UPS 4f spectra of the same compounds.<sup>12</sup> That is, the conduction-band half width and the bare 4f level are D =2.8 eV and  $\varepsilon_f$ =-1.05 eV, respectively. The 4*f*-level hybridization had been determined as  $\Delta = 116$  meV for x  $\leq 0.1$  and  $\Delta = 108$  meV for  $x \geq 0.2$ , corresponding to a reduction in the effective spin exchange coupling  $J=\Delta^2[1/$  $(-\varepsilon_f) + 1/(\varepsilon_f + U_{ff})$ ]. The XPS 3d spectra involve the additional interaction parameters  $U_{df}$  and  $U_{ff}$ , see above. Since the intra-4*f*-level repulsion  $U_{ff}$  will not significantly depend on the Au concentration x, we have chosen a constant value of  $U_{ff}=10.0 \text{ eV} \ge |\epsilon_f|$  (within NCA it was assumed to be infinite for simplicity, in order to suppress 4f double occupancy). This leaves the core hole potential  $U_{df}$  as the only adjustable parameter. It determines the final-state energy levels and thus the position of the core-level peaks relative to each other according to Eq. (1). To fit the theoretical spectra to the experimental results we added to the GS spectra a background, accounting phenomenologically for inelastic scattering of the electrons during the XPS process, and convoluted the theoretical spectra with a Gaussian, accounting for the experimental resolution. The fits yield  $U_{df}$ =12.5 eV for  $x \le 0.1$  and  $U_{df} = 13.5$  eV for  $x \ge 0.2$ . Figure 2 shows the GS spectra, fitted to the experimental data, and the lower panel of Fig. 3 shows a comparison of the unbroadened GS spectra below and above  $x_c$ . The agreement between theory and experiment is remarkable: both the relative peak intensities and the shape of the spectra are well reproduced by only a single fit parameter.



FIG. 2. (Color online) Comparison of the experimental XPS data and the theoretical spectra calculated from the GS theory (Refs. 19 and 20). The dashed lines represent the inelastic background that was added phenomenologically to the theoretical spectra.

This allows for a detailed interpretation of the experimental data. A reduction in  $\Delta$  leads to a significant change in the relative weights of the peaks, as already shown by Gunnarsson and Schönhammer in Refs. 19 and 20: for  $\Delta$ =108 meV the  $f^2$  peak is diminished and the  $f^0$  peak has almost vanished. Since the weight of the  $f^0$  configuration is a measure of the weight of the Kondo resonance (see above), this means for the present  $CeCu_{6-x}Au_x$  system that the spectral weight of the Kondo resonance drops abruptly as x is changed from  $x \le x_c$  to  $x > x_c$ , in full agreement with the findings of UPS in Ref. 12. In the left upper panel of Fig. 3 we show the weight of the  $f^0$  configuration as extracted from the GS calculations for different x. It shows a significant step at  $x = x_c$ . As a consequence of this collapse of the Kondo resonance in the 4f spectrum above the Fermi level, one expects that the spectral weight below  $E_F$  and, hence, the f-level occupation number  $n_f$  are increased for  $x > x_c$  by roughly the Kondo spectral weight. As an immediate further consequence the 3d electron binding energy should decrease because the enhancement of  $n_f$  increases the attractive interaction exerted by the 4f electrons on the 3d core hole.

Both expectations can be checked from the XPS spectra. First, the decrease in 3*d* binding energy is indeed seen in the experimental spectra (Fig. 1) and in the fitted theoretical spectra (Fig. 3, lower panel) as the upward shift of the  $f^1$  and  $f^2$  peaks for  $x \ge 0.2$  as compared to  $x \le 0.1$ . It is also reflected



FIG. 3. (Color online) The upper left panel shows the weight of the  $f^0$  configuration as determined from the GS fit. It represents a measure of the weight of the Kondo-resonance in the 4f spectrum (see text). The upper right panel shows the 4f-occupation numbers. In the lower panel the calculated unbroadened GS spectra are shown for CeCu<sub>6-x</sub>Au<sub>x</sub> compounds below and above the quantum critical concentration  $x_c$ . The parameter values used in the calculations are given in the text. The vertical lines mark the  $f^0$ ,  $f^1$ , and  $f^2$ peaks and the SO shifting for  $x < x_c$ .

by the increase in the fitted values of the effective core-level potential  $U_{df}$ . Second, the probability for the 4*f* occupancy in the ground state can be extracted from the XPS spectra.<sup>15,16</sup> It is roughly given by

$$n_f = 1 - \frac{I(f^0)}{I(f^0) + I(f^1) + 2I(f^2)},$$
(2)

where  $I(f^n)$  is the (un-normalized) spectral weight of the  $f^n$  peak, n=0, 1, 2, after subtraction of the inelastic background. Determining  $n_f$  directly from the experimental data is not accurate since the  $f^n$  peaks overlap and the determination of the inelastic background of the experimental spectra is not trivial. Instead we extract the  $n_f$  numbers from the fitted GS calculations. These include also small deviations from Eq. (2) which arise from the dynamics of the XPS process.<sup>20,22</sup> In the right upper panel of Fig. 3 we compare the  $n_f$  values obtained from the GS calculations and from NCA calculations.

tions of Ref. 12. Since  $n_f$  increases with temperature,<sup>26</sup> it is important to note that the NCA values have been calculated at the experimental temperature of T=15 K. Although the absolute values of  $n_f$  differ for the two calculations probably due to the use of a infinite  $U_{ff}$  within NCA—the main features are the same: both data sets exhibit a significant abrupt step of  $n_f$  at the critical concentration  $x_c$  for the QPT in CeCu<sub>6-x</sub>Au<sub>x</sub>, as expected from the Kondo resonance collapse observed for x > 0.1.

## **V. CONCLUSION**

We have presented data from XPS measurements at elevated temperature on  $\text{CeCu}_{6-x}\text{Au}_x$  compounds with five different Au concentrations x across the critical concentration  $x_c=0.1$  of the magnetic quantum phase transition. With increasing x the 3d core spectra exhibit near  $x=x_c$  an abrupt change with respect to three different features, namely, (1) a collapse of the  $f^0$  resonance, signaling a sudden decrease in Kondo screening, (2) a steplike increase in the 4f occupation number, and (3) a sudden increase in the core-hole attraction, evidenced by a shift of the  $f^1$  and  $f^2$  peaks toward smaller binding energy. The first feature is in complete agreement with findings of direct UPS measurements of the Kondo resonance in the Ce 4f spectrum. The features (2) and (3) are explained in a natural way as direct consequences of the Kondo resonance collapse. Therefore, the present measurements strongly support that the Kondo resonance collapse observed at the QPT by UPS (Ref. 12) is not a surface effect but an effect of the bulk CeCu<sub>6-r</sub>Au<sub>r</sub> compounds since it is seen by the bulk-sensitive XPS as well. The analysis of the spectra using the Gunnarson-Schönhammer theory,<sup>19,20</sup> supported by NCA calculations, further indicates that the spectra taken at temperatures above the single-impurity Kondo temperature  $T_K$  and above the lattice coherence and Néel temperatures,  $T_{\rm coh}$  and  $T_N$ , are well described by the singleimpurity Anderson model, with the spin exchange coupling Jrenormalized only by RKKY interactions with the surrounding Ce impurity moments.<sup>12</sup> Thus, our results indicate experimentally that the HM model<sup>7,8</sup> is insufficient to describe the QPT in  $CeCu_{6-x}Au_x$  and that at the lowest temperatures the system should follow the local quantum critical scenario,<sup>9,10</sup> as argued in detail in Ref. 12.

#### ACKNOWLEDGMENTS

We would like to thank O. Gunnarsson for fruitful discussions and for the allocation of his code. This work was supported by the Deutsche Forschungsgemeinschaft through Grants No. Re 1469/4-3/4 (M.K. and F.R.), No. FOR 960 (H.v.L.), and No. SFB 608 (J.K.).

- <sup>1</sup>H. von Löhneysen, A. Schröder, and O. Stockert, J. Alloys Compd. **303-304**, 480 (2000).
- <sup>2</sup>O. Stockert, H. von Löhneysen, A. Rosch, N. Pyka, and M. Loewenhaupt, Phys. Rev. Lett. **80**, 5627 (1998).
- <sup>3</sup>A. Schröder, G. Aeppli, R. Coldea, M. Adams, O. Stockert, H. von Löhneysen, E. Bucher, R. Ramazashvili, and P. Coleman, Nature (London) **407**, 351 (2000).
- <sup>4</sup>H. von Löhneysen, A. Neubert, T. Pietrus, A. Schröder, O. Stockert, U. Tutsch, M. Löwenhaupt, A. Rosch, and P. Wölfle, Eur. Phys. J. B **5**, 447 (1998).
- <sup>5</sup>H. von Löhneysen, H. Bartolf, S. Drotziger, C. Pfleiderer, O. Stockert, D. Souptel, W. Löser, and G. Behr, J. Alloys Compd. **408-412**, 9 (2006).
- <sup>6</sup>P. Gegenwart, Q. Si, and F. Steglich, Nat. Phys. 4, 186 (2008).
- <sup>7</sup>J. A. Hertz, Phys. Rev. B **14**, 1165 (1976).
- <sup>8</sup>A. J. Millis, Phys. Rev. B **48**, 7183 (1993).
- <sup>9</sup>Q. Si, S. Rabello, and J. L. Smith, Nature (London) **413**, 804 (2001).
- <sup>10</sup>P. Coleman, C. Pépin, Q. Si, and R. Ramazashvili, J. Phys.: Condens. Matter **13**, R723 (2001).
- <sup>11</sup>A. Rosch, A. Schröder, O. Stockert, and H. von Löhneysen, Phys. Rev. Lett. **79**, 159 (1997).
- <sup>12</sup>M. Klein, A. Nuber, F. Reinert, J. Kroha, O. Stockert, and H. von Löhneysen, Phys. Rev. Lett. **101**, 266404 (2008).
- <sup>13</sup>D. Ehm, S. Hüfner, F. Reinert, J. Kroha, P. Wölfle, O. Stockert, C. Geibel, and H. von Löhneysen, Phys. Rev. B **76**, 045117 (2007).
- <sup>14</sup>J. W. Allen, J. Phys. Soc. Jpn. 74, 34 (2005).
- <sup>15</sup>F. Reinert, R. Claessen, G. Nicolay, D. Ehm, S. Hüfner, W. P.

Ellis, G. H. Gweon, J. W. Allen, B. Kindler, and W. Aßmus, Phys. Rev. B **63**, 197102 (2001).

- <sup>16</sup>S. Schmidt, N. Weber, H. J. Elmers, F. Forster, F. Reinert, S. Hüfner, M. Escher, M. Merkel, B. Krömker, and D. Funnemann, Phys. Rev. B **72**, 064429 (2005).
- <sup>17</sup>S. Hüfner, *Photoelectron Spectroscopy* (Springer-Verlag, Berlin, Heidelberg, New York, 1994).
- <sup>18</sup>K. Grube, W. H. Fietz, U. Tutsch, O. Stockert, and H. von Löhneysen, Phys. Rev. B **60**, 11947 (1999).
- <sup>19</sup>O. Gunnarsson and K. Schönhammer, Phys. Rev. B 28, 4315 (1983).
- <sup>20</sup>O. Gunnarsson, K. Schönhammer, J. C. Fuggle, F. U. Hillebrecht, J. M. Esteva, R. C. Karnatak, and B. Hillebrand, Phys. Rev. B 28, 7330 (1983).
- <sup>21</sup>D. Ehm, F. Reinert, G. Nicolay, S. Schmidt, S. Hüfner, R. Claessen, V. Eyert, and C. Geibel, Phys. Rev. B 64, 235104 (2001).
- <sup>22</sup>J. C. Fuggle, F. U. Hillebrecht, Z. Zolnierek, R. Lässer, C. Freiburg, O. Gunnarsson, and K. Schönhammer, Phys. Rev. B 27, 7330 (1983).
- <sup>23</sup>J. W. Allen, S. J. Oh, O. Gunnarsson, K. Schönhammer, M. B. Maple, M. S. Torikachvili, and I. Lindau, Adv. Phys. **35**, 275 (1986).
- <sup>24</sup>N. Witkowski, F. Bertran, and D. Malterre, Phys. Rev. B 56, 15040 (1997).
- <sup>25</sup>N. Witkowski, F. Bertran, and D. Malterre, J. Electron Spectrosc. Relat. Phenom. **117-118**, 371 (2001).
- <sup>26</sup>N. E. Bickers, D. L. Cox, and J. W. Wilkins, Phys. Rev. B 36, 2036 (1987).